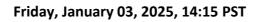
3-day Space Weather Conditions (SUPARCO)





Radio Blackouts			Solar Radiation Storms			Geomagnetic Storms		
-24 Hr	Current	Predicted	-24 Hr	Current	Predicted	-24 Hr	Current	Predicted
R1	R0	R0 – R1	SO	S 0	S 0	G1	G0	G0

		LOC	AL CURREN	IT IONOSPHERIC CONDIT	IONS (SON)				
DATE	3-Jan-25 (noon)			4-Jan-25		5-Jan-25 (noon)			
foF2	14.5 MHz			14.2		14.1 MHz			
h′F2	318 km			290		287 km			
TEC	68 TECU			63 TE		60 TECU			
	Maximum Usa	ble Frequency	(MUF) and	d Optimum Traffic Fre	quency (FOT) fo	r various dista	inces		
Distance (km)	100	200	400	600	800	1000	1500	3000	
MUF (MHz) for 3	14.8	15.2	17.2	2 20.0	23.1	26.3	29.9	33.9	
days (03 Jan – 05	14.4	14.9	16.9		22.9	26.2	29.7	33.7	
Jan)	14.3	14.7	16.7		22.6	26.0	29.4	33.4	
FOT (MHz) for 3	12.5	12.9	14.6	5 17.0	19.6	22.4	25.5	28.8	
days (03 Jan – 05	12.2	12.6	14.4	4 16.8	19.5	22.2	25.2	28.7	
Jan)	12.0	12.4	14.1		19.2	21.9	24.9	28.4	
Local ionospheric condit	ions are slightly e	enhanced as com							
			LOCAL	GEOMAGNETIC CONDITIO	NS				
K-index	2 (Quiet)			Quiet geomagnetic activity is expected.		. Quiet geo	Quiet geomagnetic activity is expected.		
F (SON/ISB)	45675/50515 nT			45682±10 /50520±20 nT		456	45682±10/50520±20 nT		
The local geomagnetic field	d is quiet at the mo	oment.							
				SOLAR CONDITIONS					
SN	173		162 (SSN-predicted)		1	153 (SSN-predicted)			
F 10.7	212 sfu			207 sfu			187 sfu		
Vsw	448.7 km/s (Varied in the past 12			Low to moderate le	d Low to m	Low to moderate levels of solar wind			
VSW	hrs between 423 & 552 km/s)			speed may		speed may prevail.			
Solar flares	C3.2 (max. flare in the past			Low to moderat	Low	Low to moderate levels of solar			
Solar flares	(M1, 1740 UT)			activity is expected.		ā	activity is expected.		
IMF	IMF +14.73 nT (varied in the past 12 hrs between +7.94 nT & +15.72 nT)								
Bt									
				Expected to vary be			d to vary betwe	•	
Bz	+11.82 nT (varied in the past 12 hrs			negative sectors.		a	and negative sectors.		
D2	between +0.45 nT & +8.16 nT)								

Solar conditions are at low to moderate levels with background X-ray flux at C-class levels.							
Daily Sun: 3 January 2025							
There are two active regions AR3945 and AR3947 present on the Sun capable of producing strong solar flares. 02 Coronal Holes (CHs) are detected on the solar disk.	NSO/GONG 677nm Learmonth (AUS) 3948 3938 3947 3944 3943 3944 3939 3935 2025-01-03 01:34Z COES-16 SUVI (1/3/2025 & 01:08 UTC)						
2-Day Conditions							
 Solar activity is expected to be at low to moderate levels. In case of M/X-class solar flares, minor level radio blackouts Low to moderate levels of solar wind speed is expected due Quiet geomagnetic activity is expected to prevail. 	•						

Slightly enhanced ionospheric conditions are expected for the next 2 days. It is advised to use the frequency ranges mentioned in the ionospheric section.

Credits:

Solar conditions courtesy to SOHO, DSCOVR and GOES-16 missions. NOAA SWPC is acknowledged for solar radio flux conditions. Korean Space Weather Centre is acknowledged for solar disk and coronal hole images.

Sonmiani (SON): 25.2° N, 66.75° E Islamabad (ISB): 33.7° N, 73.13° E

<u>ANNEXURE</u>

	DEFINITIONS OF TERMINOLOGIES USED IN THIS SUMMARY
foF2	Maximum frequency of F2-layer of the ionosphere
h′F2	Virtual height of the F2-layer
MUF	Maximum usable frequency for 3000 km
K-index	Local index defining geomagnetic conditions
Declination	Planetary A index defining geomagnetic conditions, predicted value during geomagnetic unsettled Conditions
F	Magnitude of the total geomagnetic field vector (unit in nano Teslas)
SON, difference	Sonmiani Geomagnetic Observatory mean value, <u>difference limit</u> from night time value of quiet conditions: 25-30 nT, max: 260 nT
ISB	Islamabad Geomagnetic Observatory mean value
SN	Relative sunspot numbers
Vsw	Solar Wind Speed (km/s)
F10.7	Solar radio flux at 2.8 GHz (10.7 cm wavelength)
sfu	Solar flux unit (defines the solar radio 10.7 cm flux)
Solar Flare	Could be B, C, M and X depending upon the intensity of x-rays being emitted (each type has further 10 classes based on amount of energy released by the flare)
IMF	Interplanetary magnetic field (the source of which is the Sun)
Bt	Total IMF (unit in Nano Teslas)
Bz	Vertical component of IMF (could be north/upward/positive or south/downward/negative) (unit in nano Teslas)
AR	Active Regions on the sun currently in view
CME	Coronal Mass Ejection
СН	Coronal Hole
KASI	Korean Astronomy & Space Science Institute
SWFs	Short-wave fadeouts, caused by M/X class flares on the day lit side of the hemisphere absorbing lower Frequencies and hampering HF communication.
SSN-predicted	Smooth Sunspot Number-it is an estimated value using a mathematical relation to forecast it.

<u>RSG SCALES</u>

	Radio Blackouts							
Minor	Moderate	oderate Strong Se		Extreme				
R1	R2	R3	R4	R5				
	Solar Radiation Storms							
Minor	Moderate	te Strong Se		Extreme				
S1	S2	S3	S4	S5				
	Geomegnatic Storms							
Minor	Moderate	Strong	Severe	Extreme				
G1	G2	G3	G4	G5				