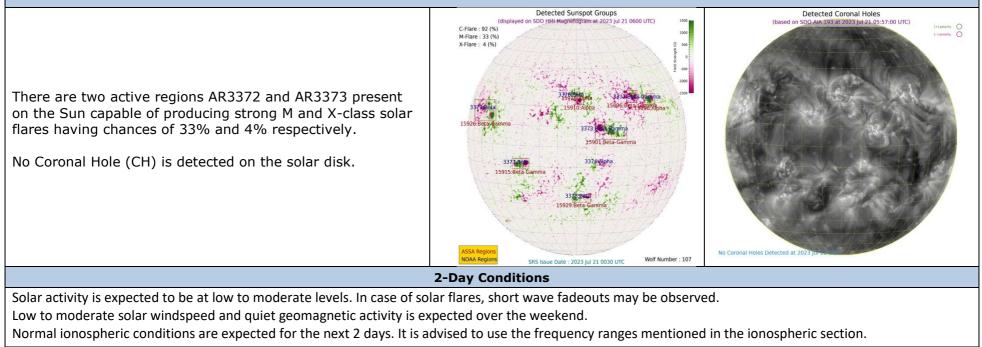
3-day Space Weather Conditions (SUPARCO)

Friday, July 21, 2023, 12:13 PST



		LOCA	AL CURREN	IT IONOSPHERIC CONDI	TIONS (SON)				
DATE	21-Jul-23 (noon)			22-Jul-23		23-Jul-23 (noon)			
foF2	10.4 MHz			10.2		10.0 MHz			
h′F2	310 km			290 km			278 km		
TEC	47 TECU			45 T		42 TECU			
	Maximum Usa	able Frequency ((MUF) and	d Optimum Traffic Fre	equency (FOT) for	various dista	inces		
istance (km)	100	200	400	600	800	1000	1500	3000	
UF (MHz) for 3	10.5	10.9	12.3	3 14.0	16.3	18.3	22.4	28.0	
ays (21 Jul – 23	10.3	10.7 12		1 13.8	15.9	18.0	21.9	27.7	
Í)	10.1	10.5	11.9	9 13.5	15.4	17.8	21.5	27.3	
DT (MHz) for 3	8.9	9.3	10.3	3 11.9	14.8	15.5	19.9	23.8	
nys (21 Jul – 23	8.8	9.1	10.2	2 11.7	14.5	15.3	18.6	23.5	
l)			10.1		13.1	15.1	18.3	23.2	
cal ionospheric condition	tions are normal a	as compared to th	e predicte	d monthly median MU					
			LOCAL	GEOMAGNETIC CONDITIC	NS				
K-index	2 (Quiet)			Quiet geomagnetic activity is expected.		Quiet geo	Quiet geomagnetic activity is expected		
F (SON/ISB)	45525/50035 nT			45538±10/5	45	45538±10/50045±20 nT			
ne local geomagnetic fiel	d is Quiet at the mo	oment.							
				SOLAR CONDITIONS					
SN	131			128 (SSN-predicted)		1	125 (SSN-predicted)		
F 10.7	184 sfu			182		180 sfu			
Vsw	412.3 km/s (Varied in the past 12 hrs between 346 & 487 km/s)			Low to moderate levels of solar		Low	Low to moderate levels of sola windspeed may prevail.		
				windspeed may prevail.		WI			
Solar flares	C1.5 (max. flare in the past 24 hrs: M2, 1920 UT)			Moderate levels of solar activity		Low	Low to moderate levels of sola activity expected.		
				expec					
IMF	+5.2 nT (va	ried in the past	12 hrs						
Bt	between +6.9 nT & +11.7 nT)								
DL				Expected to vary be		Expected to vary between positiv and negative sectors.			
D -	+2.8 nT (varied in the past 12 hrs between -7.0 nT & +4.7 nT) derate levels with background X-ray flux at M-clas			negative	a				
Bz									

Daily Sun: 21 July 2023



<u>For information on radio blackout levels, please follow the link:</u> http://www.swpc.noaa.gov/noaa-scales-explanation

Acknowledgements:

<u>Images source</u>: Solar Dynamics Observatory-SDO both images showing the Solar disk and Coronal Holes have been processed at SUPARCO using Automatic Solar Synoptic Analyzer (ASSA), developed jointly by the Korean Space Weather Centre of the Radio Research Agency (RRA) & Space Environment Laboratory (SE Lab).

<u>Data sources</u>: The planetary indices and solar data are taken from the URLs below:

http://<u>www.spaceweather.go.kr</u> http://<u>www.sws.bom.gov.au</u> http://<u>www.solarmonitor.org</u>

Sonmiani (SON): 25.2° N, 66.75° E Islamabad (ISB): 33.7° N, 73.13° E

ANNEXURE

	DEFINITIONS OF TERMINOLOGIES USED IN THIS SUMMARY					
foF2	Maximum frequency of F2-layer of the ionosphere					
h′F2	Virtual height of the F2-layer					
MUF	Maximum usable frequency for 3000 km					
K-index	Local index defining geomagnetic conditions					
Declination	Planetary A index defining geomagnetic conditions, predicted value during geomagnetic unsettled Conditions					
F	Magnitude of the total geomagnetic field vector (unit in nano Teslas)					
SON, difference	Sonmiani Geomagnetic Observatory mean value, <u>difference limit</u> from night time value of quiet conditions: 25-30 nT, max: 260 nT					
ISB	Islamabad Geomagnetic Observatory mean value					
SN	Relative sunspot numbers					
Vsw	Solar Wind Speed (km/s)					
F10.7	Solar radio flux at 2.8 GHz (10.7 cm wavelength)					
sfu	Solar flux unit (defines the solar radio 10.7 cm flux)					
Solar Flare	Could be B, C, M and X depending upon the intensity of x-rays being emitted (each type has further 10 classes based on amount of energy released by the flare)					
IMF	Interplanetary magnetic field (the source of which is the Sun)					
Bt	Total IMF (unit in Nano Teslas)					
Bz	Vertical component of IMF (could be north/upward/positive or south/downward/negative) (unit in nano Teslas)					
AR	Active Regions on the sun currently in view					
CME	Coronal Mass Ejection					
СН	Coronal Hole					
KASI	Korean Astronomy & Space Science Institute					
SWFs	Short-wave fadeouts, caused by M/X class flares on the daylit side of the hemisphere absorbing lower Frequencies and hampering HF communication.					
SSN-predicted	Smooth Sunspot Number-it is an estimated value using a mathematical relation to forecast it.					