3-day Space Weather Conditions (SUPARCO)

Friday, April 28, 2023, 12:01 PST



		LOC	CAL CURRENT ION	IOSPHERIC CONDI	TIONS (SON)			
DATE	28-	Apr-23 (noon)		29-Apr-23	3 (noon)		30-Apr-23 (noc	on)
foF2	12.2 MHz			12.0 MHz			11.8 MHz	
h′F2	340 km			337 km			332 km	
TEC	52 TECU			50 TECU			49 TECU	
	Maximum Us	able Frequency	(MUF) and Opt	imum Traffic Fr	equency (FOT)	for various dist	ances	
Distance (km)	100	200	400	600	800	1000	1500	3000
MUF (MHz) for 3	12.4	12.7	14.0	15.8	17.8	20.2	25.2	30.6
days (28 Ápr – 30	12.2	12.5	13.8	15.6	17.7	19.8	25.0	30.2
Apr)	12.0	12.4	13.6	15.4	17.5	19.7	24.8	30.0
FOT (MHz) for 3	10.5	10.8	11.9	13.4	15.1	17.2	21.4	26.0
days (28 Ápr – 30	10.4	10.6	11.7	13.3	15.0	16.8	21.3	25.7
Apr)	10.2	10.5	11.6	13.1	14.9	16.7	21.1	25.5

Local ionospheric conditions are slightly enhanced as compared to the predicted monthly median MUF.

LOCAL GEOMAGNETIC CONDITIONS				
K-index	2 (Quiet)	Quiet geomagnetic activity is expected.	Quiet geomagnetic activity is expected.	
F (SON/ISB)	45515/50025 nT	45528±10 /50035±20 nT	45528±10/50035±20 nT	

The local geomagnetic field is disturbed at the moment.

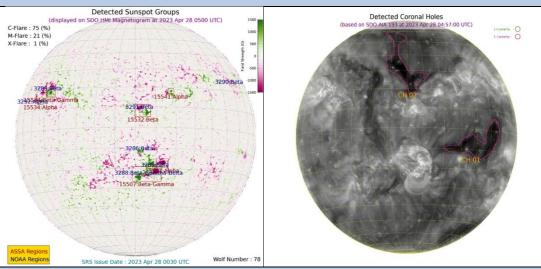
		SOLAR CONDITIONS		
SN	136	134 (SSN-predicted)	130 (SSN-predicted)	
F 10.7	141 sfu	137 sfu	132 sfu	
V _{SW}	566.3 km/s (Varied in the past 12 hrs between 527 & 672 km/s)	Moderate to high levels of solar wind speed may prevail.	Moderate levels of solar wind speed may prevail.	
Solar flares	C1.6 (max. flare in the past 24 hrs: M1, 1114 UT)	Low to moderate levels of solar activity expected.	Low levels of solar activity expected.	
IMF Bt	+5.1 nT (varied in the past 12 hrs between +4.8 nT & +5.6 nT)	Expected to vary between positive and	Expected to vary between positive and	
Bz	+2.1 nT (varied in the past 12 hrs between -4.6 nT & +4.6 nT)	negative sectors.	negative sectors.	

Solar conditions are at low to moderate levels with background X-ray flux at C-class levels.

Daily Sun: 28 April 2023

There is one active region AR3288 present on the Sun capable of producing strong M and X-class solar flares having chances of 21% and 1% respectively.

02 Coronal Holes (CHs) are detected on the solar disk.



2-Day Conditions

Solar activity is expected to be at low to moderate levels. In case of solar flares, short wave fadeouts may be observed.

Moderate to elevated solar windspeed is expected due to the presence of coronal holes.

Quiet geomagnetic activity is expected over the weekend.

Slightly enhanced ionospheric conditions are expected for the next 2 days. It is advised to use the frequency ranges mentioned in the ionospheric section.

For information on radio blackout levels, please follow the link:

http://www.swpc.noaa.gov/noaa-scales-explanation

<u>Acknowledgements:</u>

<u>Images source</u>: Solar Dynamics Observatory-SDO both images showing the Solar disk and Coronal Holes have been processed at SUPARCO using Automatic Solar Synoptic Analyzer (ASSA), developed jointly by the Korean Space Weather Centre of the Radio Research Agency (RRA) & Space Environment Laboratory (SE Lab).

<u>Data sources</u>: The planetary indices and solar data are taken from the URLs below:

http://www.spaceweather.go.kr http://www.sws.bom.gov.au http://www.solarmonitor.org

Sonmiani (SON): 25.2º N, 66.75º E Islamabad (ISB): 33.7º N, 73.13º E

ANNEXURE

	DEFINITIONS OF TERMINOLOGIES USED IN THIS SUMMARY			
foF2	Maximum frequency of F2-layer of the ionosphere			
h′F2	Virtual height of the F2-layer			
MUF	Maximum usable frequency for 3000 km			
K-index	Local index defining geomagnetic conditions			
Declination	Planetary A index defining geomagnetic conditions, predicted value during geomagnetic unsettled Conditions			
F	Magnitude of the total geomagnetic field vector (unit in nano Teslas)			
SON, difference	Sonmiani Geomagnetic Observatory mean value, <u>difference limit</u> from night time value of quiet conditions: 25-30 nT, max: 260 nT			
ISB	Islamabad Geomagnetic Observatory mean value			
SN	Relative sunspot numbers			
Vsw	Solar Wind Speed (km/s)			
F10.7	Solar radio flux at 2.8 GHz (10.7 cm wavelength)			
sfu	Solar flux unit (defines the solar radio 10.7 cm flux)			
Solar Flare	Could be B, C, M and X depending upon the intensity of x-rays being emitted (each type has further 10 classes based on amount of energy released by the flare)			
IMF	Interplanetary magnetic field (the source of which is the Sun)			
Bt	Total IMF (unit in Nano Teslas)			
Bz	Vertical component of IMF (could be north/upward/positive or south/downward/negative) (unit in nano Teslas)			
AR	Active Regions on the sun currently in view			
СМЕ	Coronal Mass Ejection			
СН	Coronal Hole			
KASI	Korean Astronomy & Space Science Institute			
SWFs	Short-wave fadeouts, caused by M/X class flares on the daylit side of the hemisphere absorbing lower Frequencies and hampering HF communication.			
SSN-predicted	Smooth Sunspot Number-it is an estimated value using a mathematical relation to forecast it.			