3-day Space Weather Conditions (SUPARCO)

Friday, December 16, 2022, 12:47 PST



| | | LOC | CAL CURREN | NT IONOSPHERIC COND | ITIONS (SON) | | | |
|-------------------------------|---|-------------------|------------|----------------------|----------------|---|-------|------|
| DATE | 16-Dec-22(noon) | | | 17-Dec-2 | | 18-Dec-22 (noon) | | |
| foF2 | 9.9 MHz | | | 10.2 | | 9.1 MHz | | |
| h′F2 | 300 km | | | 315 | | 290 km | | |
| TEC | 44 TECU | | | 48 T | | 38 TECU | | |
| | Maximum Us | able Frequency | (MUF) an | d Optimum Traffic F | requency (FOT) | for various dist | ances | |
| Distance (km) | 100 | 200 | 400 | 600 | 800 | 1000 | 1500 | 3000 |
| MUF (MHz) for 3 | 10.1 | 10.4 | 11.6 | 5 13.2 | 15.1 | 17.0 | 21.4 | 26.6 |
| lays (16 Dec - 18 | 10.4 | 10.7 | 11.9 | | 15.2 | 17.1 | 21.6 | 26.9 |
| Dec) | 9.3 | 9.7 | 10.8 | 3 12.4 | 14.2 | 16.0 | 20.2 | 24.9 |
| OT (MHz) for 3 | 8.6 | 8.8 | 9.9 | 11.2 | 12.8 | 14.5 | 18.2 | 22.6 |
| lays (16 Dec - 18 | 8.8 | | | . 11.4 | 12.9 | 14.6 | 18.4 | 22.9 |
| Dec) | | | 9.2 | 10.5 | 12.1 | 13.6 | 17.2 | 21.2 |
| ocal ionospheric condit. | ions are normal a | as compared to th | e predicte | d monthly median MU | F. | | | |
| | | | LOCA | L GEOMAGNETIC CONDIT | IONS | | | |
| K-index | 1 | | | Quiet geomagnetic a | d. Quiet ge | Quiet geomagnetic activity is expected | | |
| F (SON/ISB) | 45533/50043 nT | | | 45545±10/5 | 4! | 45545±10/50050±20 nT | | |
| The local geomagnetic fiel | d is quiet at the mo | oment. | | | | | | |
| | | | | SOLAR CONDITIONS | | | | |
| SN | 140 | | | 157 (SSN- | | 133 (SSN-predicted) | | |
| F 10.7 | 166 sfu | | | 175 | | 142 sfu | | |
| Vsw | 315.3 km/s (Varied in the past 12 hrs between 316 & 546 km/s) | | | Clightly onhoncod l | nd Mo | Moderate levels of solar wind speed may prevail. | | |
| | | | | Slightly enhanced I | MO MO | | | |
| | | | | speed may prevail. | | | | |
| Solar flares | M1.1 (max. flare in the past 24 hrs: M5, 2240 UT) | | | Moderate level | y Lov | Low to moderate levels of sola activity expected. | | |
| Solar lidres | | | | expe | | | | |
| IMF | +5.03 nT (va | aried in the past | 12 hrs | - | | | | |
| Bt | between +4.2 nT & +6.5 nT) | | | | | | | |
| DL | | | | Expected to vary be | and Expect | Expected to remain in the negative | | |
| R-7 | +0.87 nT (varied in the past 12 hrs | | | negative | | sectors. | | |
| Bz between -4.2 nT & +4.6 nT) | | | | | | | | |

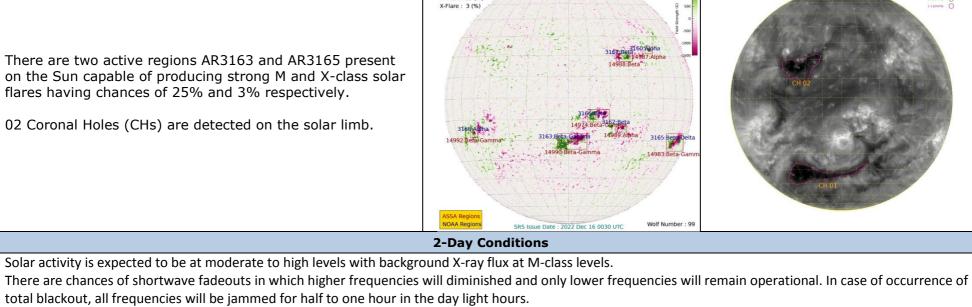
Daily Sun: 16 December 2022

C-Flare : 85 (%)

M-Flare : 25 (%

Detected Sunspot Groups ed on SDO HMI Magnetogram at 2022 Dec 16 0600 UTC)

Detected Coronal Holes SDO AIA 193 at 2022 Dec 16 05:57:00 UTC



Light to slightly elevated solar wind speed is expected to prevail due to the presence of coronal hole.

Quiet geomagnetic activity is expected to prevail over the weekend.

Normal to slightly enhanced ionospheric conditions are expected for the next 2 days. It is advised to use the frequency ranges mentioned in the ionospheric section.

<u>For information on radio blackout levels, please follow the link:</u> http://www.swpc.noaa.gov/noaa-scales-explanation

Acknowledgements:

Images source: Solar Dynamics Observatory-SDO both images showing the Solar disk and Coronal Holes have been processed at SUPARCO using Automatic Solar Synoptic Analyzer (ASSA), developed jointly by the Korean Space Weather Centre of the Radio Research Agency (RRA) & Space Environment Laboratory (SE Lab).

<u>Data sources</u>: The planetary indices and solar data are taken from the URLs below:

http://<u>www.spaceweather.go.kr</u> http://www.sws.bom.gov.au

http://www.solarmonitor.org

<u>ANNEXURE</u>

| DEFINITIONS OF TERMINOLOGIES USED IN THIS SUMMARY | | | | | | |
|---|--|--|--|--|--|--|
| foF2 | Maximum frequency of F2-layer of the ionosphere | | | | | |
| h′F2 | Virtual height of the F2-layer | | | | | |
| MUF | Maximum usable frequency for 3000 km | | | | | |
| K-index | Local index defining geomagnetic conditions | | | | | |
| Declination | Planetary A index defining geomagnetic conditions, predicted value during geomagnetic unsettled Conditions | | | | | |
| F | Magnitude of the total geomagnetic field vector (unit in nano Teslas) | | | | | |
| SON, difference | Sonmiani Geomagnetic Observatory mean value, <u>difference limit</u> from night time value of quiet conditions: 25-30 nT, max: 260 nT | | | | | |
| ISB | Islamabad Geomagnetic Observatory mean value | | | | | |
| SN | Relative sunspot numbers | | | | | |
| Vsw | Solar Wind Speed (km/s) | | | | | |
| F10.7 | Solar radio flux at 2.8 GHz (10.7 cm wavelength) | | | | | |
| sfu | Solar flux unit (defines the solar radio 10.7 cm flux) | | | | | |
| Solar Flare | Could be B, C, M and X depending upon the intensity of x-rays being emitted (each type has further 10 classes based on amount of energy released by the flare) | | | | | |
| IMF | Interplanetary magnetic field (the source of which is the Sun) | | | | | |
| Bt | Total IMF (unit in Nano Teslas) | | | | | |
| Bz | Vertical component of IMF (could be north/upward/positive or south/downward/negative) (unit in nano Teslas) | | | | | |
| AR | Active Regions on the sun currently in view | | | | | |
| CME | Coronal Mass Ejection | | | | | |
| СН | Coronal Hole | | | | | |
| KASI | Korean Astronomy & Space Science Institute | | | | | |
| SWFs | Short-wave fadeouts, caused by M/X class flares on the daylit side of the hemisphere absorbing lower Frequencies and hampering HF communication. | | | | | |
| SSN-predicted | Smooth Sunspot Number-it is an estimated value using a mathematical relation to forecast it. | | | | | |