

2-Day Space Weather Conditions (SUPARCO)

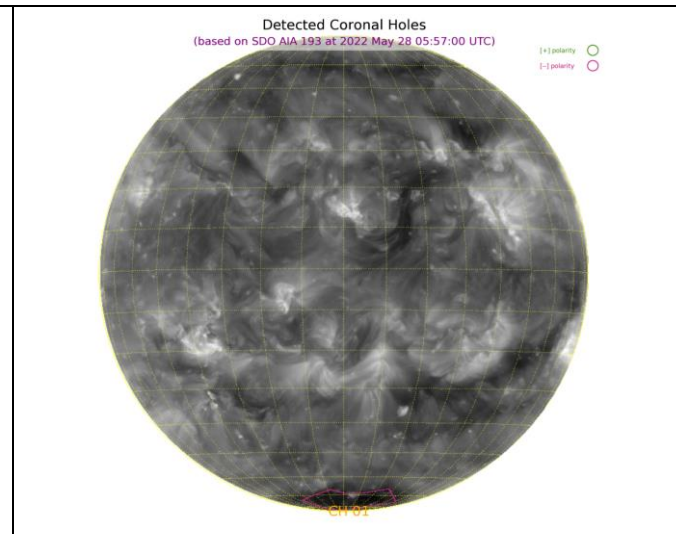
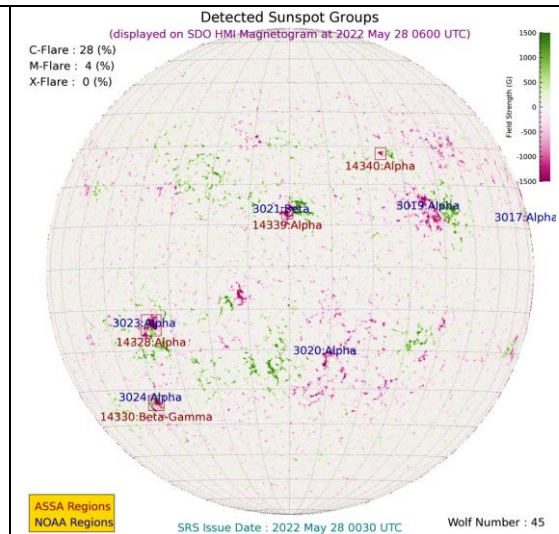
Saturday, May 28, 2022, 12:42 PST



LOCAL CURRENT IONOSPHERIC CONDITIONS (SON)								
DATE	28-May-22				29-May-22 (noon)			
foF2	10.5 MHz				10.1 MHz			
h'F2	363 km				358 km			
TEC	57 TECU				55 TECU			
Maximum Usable Frequency (MUF) and Optimum Traffic Frequency (FOT) for various distances								
Distance (Km)	100	200	400	600	800	1000	1500	3000
MUF (MHz) for 2 days(28 & 29 May)	10.6	10.9	11.9	13.3	15.0	16.8	21.2	26.2
	9.0	9.3	10.1	11.3	12.8	14.3	18.0	22.3
FOT (MHz) for 2 days(28 & 29 May)	10.2	10.5	11.5	12.9	14.6	16.4	20.8	26.0
	8.7	8.9	9.8	11.0	12.4	13.9	17.7	21.9
Local ionospheric conditions are normal as compared to predicted monthly median MUF.								
LOCAL GEOMAGNETIC CONDITIONS								
K-index	4 (Unsettled)				Quiet geomagnetic activity expected			
F(SON/ISB)	45464/50074 nT				45473±30/50079±20nT			
The local geomagnetic field is quiet at the moment.								
SOLAR CONDITIONS								
SN	69				62 (SSN-predicted)			
F 10.7	114 sfu				105 sfu			
V_{sw}	539.4 km/sec (varied in the past 12 hrs between 308 & 564 km/s)				Moderate solar wind speed is expected.			
Solar flares	B 5.6 (max. flare in the past 24 hrs: C2, 1127 UT)				Low to moderate levels of solar activity expected.			
IMF Bt Bz	6.8 nT (varied in the past 12 hrs between 7.1 & 9.3 nT) 4.5 nT (varied in the past 12 hrs between -7.6 & 4.4 nT)				Expected to vary between positive and negative sectors.			
Solar conditions are at low to moderate levels with background X-ray flux at B-class levels.								
Daily Sun: 28 May 2022								

There is no active region present on the Sun capable of producing strong solar flares.

01 Coronal Hole (CH) is detected on the solar disk.



1-Day Conditions

Solar activity is expected to be at low to moderate levels.

Moderate solar wind speed is expected due to geo-effective coronal hole.

Unsettled geomagnetic conditions are prevalent which are expected to subside within the next 24 hours.

Local Ionospheric conditions are normal. It is advised to use the frequency ranges mentioned in the ionospheric section.

For information on radio blackout levels, please follow the link:

<http://www.swpc.noaa.gov/noaa-scales-explanation>

Acknowledgements:

Images source: Solar Dynamics Observatory-SDO) Both images showing the Solar disk and Coronal Holes have been processed at SUPARCO using Automatic Solar Synoptic Analyzer (ASSA), developed jointly by the Korean Space Weather Centre of the Radio Research Agency (RRA) & Space Environment Laboratory (SELab).

Data sources: The planetary indices and solar data are taken from the URLs below:

<http://www.spaceweather.go.kr>

<http://www.sws.bom.gov.au>

<http://www.solarmonitor.org>

Sonmiani (SON): 25.2° N, 66.75° E

ANNEXURE

DEFINITIONS OF TERMINOLOGIES USED IN THIS SUMMARY	
foF2	Maximum frequency of F2-layer of the ionosphere
h'F2	Virtual height of the F2-layer
MUF	Maximum usable frequency for 3000 km
K-index	Local index defining geomagnetic conditions
Declination	Planetary A index defining geomagnetic conditions, predicted value during geomagnetic unsettled conditions
F	Magnitude of the total geomagnetic field vector (unit in nano Teslas)
SON, difference	Sonmiani Geomagnetic Observatory mean value, <u>difference limit</u> from night time value of quiet conditions: 25-30 nT, max: 260 nT
ISP	Islamabad Geomagnetic Observatory mean value
SN	Relative sunspot numbers
V _{sw}	Solar Wind Speed (km/s)
F10.7	Solar radio flux at 2.8 GHz (10.7 cm wavelength)
sfu	Solar flux unit (defines the solar radio 10.7 cm flux)
Solar Flare	Could be B, C, M and X depending upon the intensity of x-rays being emitted (each type has further 10 classes based on amount of energy released by the flare)
IMF	Interplanetary magnetic field (the source of which is the Sun)
B _t	Total IMF (unit in nano Teslas)
B _z	Vertical component of IMF (could be north/upward/positive or south/downward/negative) (unit in nano Teslas)
AR	Active Regions on the sun currently in view
CME	Coronal Mass Ejection
CH	Coronal Hole
KASI	Korean Astronomy & Space Science Institute
SWFs	Short-wave fadeouts, caused by M/X class flares on the daylit side of the hemisphere absorbing lower Frequencies and hampering HF communication.
SSN-predicted	Smooth Sunspot Number-it is an estimated value using a mathematical relation to forecast it.