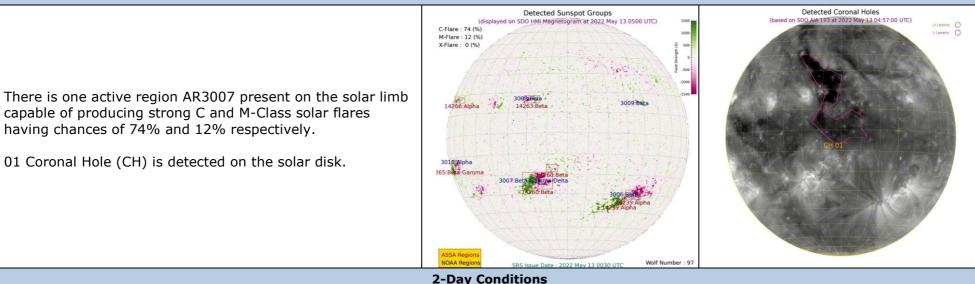
Friday, May 13, 2022, 12:09PST



DATE	13-May-22(noon)			NT IONOSPHERIC CONDITIONS (SON) 14-May-22 (noon)			15-May-22 (noon)		
foF2	10.6 MHz			10.3 MHz			10.0 MHz		
h′F2	325 km			320 km			310 km		
TEC	50 TECU			47 TECU			45 TECU		
	Maximum Us	able Frequency	(MUF) an	d Optimum Traf	ic Frequency (FO	T) for v	arious dist	ances	
istance (km)	100	200	400	600	800		1000	1500	3000
IUF (MHz) for 3	10.7	11.1	12.4	14.1	16.1		18.2	23.2	28.4
ays (13 May-15	10.4	10.7	12.0	13.7	15.7		17.8	22.8	28.0
lay)	10.1	10.4	11.7	' 13.4	15.4		17.5	22.5	27.7
OT (MHz) for 3	9.1	9.0	9.9	11.1	12.6		14.0	17.7	24.2
ays (13 May-15	8.8	8.7	9.6	10.8	12.3		13.7	17.4	23.9
lay)	8.6 8.5		9.4	10.6	12.1		13.5	17.2	23.7
ocal ionospheric condit	ions are normal a	s compared to th	e predicte	d monthly median	MUF.				
			LOCA	L GEOMAGNETIC CO	NDITIONS				
K-index	1		Quiet geomagnetic activity expected			Quiet geomagnetic activity expected			
F (SON/ISB)	45579/50089 nT			45586±10 /50092±20 nT			45586±10/50092±20 nT		
he local geomagnetic field	d is quite at the mo	ment.							
				SOLAR CONDITIO	NS				
SN	112			115 (SSN-predicted)			117 (SSN-predicted)		
F 10.7	133 sfu			138 sfu			143 sfu		
V _{sw}	331.9 km/s			Low to medarate lovels of color wind			Low to moderate levels of solar wind		
	(varied in the past 12 hrs between			Low to moderate levels of solar wind					
	298 & 382 km/s)			speed is expected.			speed is expected.		
Solar flares	C1.2 (max. flare in the past 24			Moderate to high levels of solar			Moderate levels of solar activity		
	hrs: M1 at 2019 UT)			activity expected.			expected.		
IMF	4.5 nT (vari	ed in the past 1	2 hrs						
Bt	between 4.7 nT & 7.0 nT)								
				Expected to vary between positive and negative sectors.			Expected to vary between positive an negative sectors.		
Bz	0.3 nT (varied in the past 12 hrs between -1.3 nT & 4.8 nT)								

Daily Sun: 13 May 2022



Solar activity is expected to be at moderate to high levels with background X-ray flux at C-class levels. There are slight chances of X class flares also. Low to moderate solar wind speed is expected due to coronal hole and CME impact.

Quiet to unsettled geomagnetic conditions are expected over the weekend.

Minor to moderate radio blackouts may be observed in case of solar flares.

Normal HF conditions are expected for the next 2 days. In case of high solar activity, MUF enhancement may be expected. It is advised to use the frequency ranges mentioned in the ionospheric section.

For information on radio blackout levels, please follow the link: http://www.swpc.noaa.gov/noaa-scales-explanation

having chances of 74% and 12% respectively.

01 Coronal Hole (CH) is detected on the solar disk.

Acknowledgements:

Images source: Solar Dynamics Observatory-SDO both images showing the Solar disk and Coronal Holes have been processed at SUPARCO using Automatic Solar Synoptic Analyzer (ASSA), developed jointly by the Korean Space Weather Centre of the Radio Research Agency (RRA) & Space Environment Laboratory (SELab).

Data sources: The planetary indices and solar data are taken from the URLs below:

http://www.spaceweather.go.kr http://www.sws.bom.gov.au http://www.solarmonitor.org

Sonmiani (SON): 25.2º N, 66.75º E Islamabad (ISB): 33.7º N, 73.13º E

<u>ANNEXURE</u>

	DEFINITIONS OF TERMINOLOGIES USED IN THIS SUMMARY					
foF2	Maximum frequency of F2-layer of the ionosphere					
h′F2	Virtual height of the F2-layer					
MUF	Maximum usable frequency for 3000 km					
K-index	Local index defining geomagnetic conditions					
Declination	Planetary A index defining geomagnetic conditions, predicted value during geomagnetic unsettled Conditions					
F	Magnitude of the total geomagnetic field vector (unit in nano Teslas)					
SON, difference	Sonmiani Geomagnetic Observatory mean value, <u>difference limit</u> from night time value of quiet conditions: 25-30 nT, max: 260 nT					
ISP	Islamabad Geomagnetic Observatory mean value					
SN	Relative sunspot numbers					
Vsw	Solar Wind Speed (km/s)					
F10.7	Solar radio flux at 2.8 GHz (10.7 cm wavelength)					
sfu	Solar flux unit (defines the solar radio 10.7 cm flux)					
Solar Flare	Could be B, C, M and X depending upon the intensity of x-rays being emitted (each type has further 10 classes based on amount of energy released by the flare)					
IMF	Interplanetary magnetic field (the source of which is the Sun)					
Bt	Total IMF (unit in Nano Teslas)					
Bz	Vertical component of IMF (could be north/upward/positive or south/downward/negative) (unit in nano Teslas)					
AR	Active Regions on the sun currently in view					
CME	Coronal Mass Ejection					
СН	Coronal Hole					
KASI	Korean Astronomy & Space Science Institute					
SWFs	Short-wave fadeouts, caused by M/X class flares on the daylit side of the hemisphere absorbing lower Frequencies and hampering HF communication.					
SSN-predicted	Smooth Sunspot Number-it is an estimated value using a mathematical relation to forecast it.					