3-day Space Weather Conditions (SUPARCO)



Friday, April 15, 2022, 12:11PST

LOCAL CURRENT IONOSPHERIC CONDITIONS (SON)								
DATE	15-	Apr-22(noon)		16-Apr-22	2 (noon)		17-Apr-22 (no	on)
foF2	12.9 MHz			12.7 MHz			12.5 MHz	
h′F2	293 km			290 km			288 km	
TEC	47 TECU			46 TECU			45 TECU	
	Maximum Us	able Frequency	(MUF) and	Optimum Traffic Fr	equency (FOT)	for various dist	ances	
Distance (km)	100	200	400	600	800	1000	1500	3000
MUF (MHz) for 3	13.1	13.5	15.3	17.7	20.3	23.1	29.4	35.7
days (15 Apr-17	12.9	13.3	15.1	17.5	20.1	22.9	29.2	35.5
Apr)	12.7	13.1	14.9	17.3	19.9	22.7	29.0	35.3
FOT (MHz) for 3	11.1	11.5	13.0	15.0	17.3	19.6	25.0	30.3
days (15 Ápr-17	11.0	11.3	12.8	14.9	17.1	19.5	24.8	30.1
Apr)	10.8	11.1	12.6	14.7	16.9	19.3	24.7	30.

Local ionospheric conditions are enhanced as compared to the predicted monthly median MUF.

LOCAL GEOMAGNETIC CONDITIONS				
K-index	2	Quiet geomagnetic activity expected	Quiet geomagnetic activity expected	
F (SON/ISB)	45576/50086 nT	45580±10 /50090±20 nT	45580±10/50090±20 nT	

The local geomagnetic field is quite at the moment.

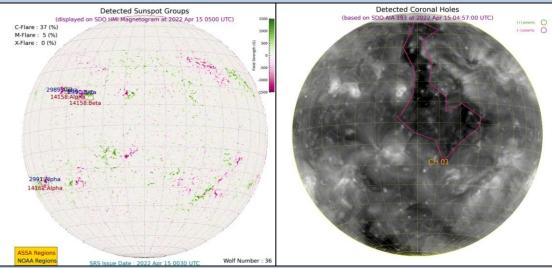
SOLAR CONDITIONS				
SN	37	36 (SSN-predicted)	35 (SSN-predicted)	
F 10.7	103 sfu	100 sfu	97 sfu	
V _{sw}	478.2 km/s (varied in the past 12 hrs between 427 & 568 km/s)	Low levels of solar wind speed expected.	Low levels of solar wind speed expected.	
Solar flares	C1.9 (max. flare in the past 24 hrs: C4 0306 UT Apr 15)	A low level of solar activity is expected.	A low level of solar activity is expected.	
IMF Bt	10.7 nT (varied in the past 12 hrs between 9.9 nT & 20.5 nT)	Expected to vary between positive and	Expected to vary between positive and	
Bz	0.1 nT (varied in the past 12 hrs between -8.7 & 10.0 nT)	negative sectors.	negative sectors.	

Solar conditions are at low levels with background X-ray flux at C-class levels.

Daily Sun: 15 April 2022

There is no active region present on the solar limb capable of producing strong solar flares.

01 Coronal Hole (CH) is detected on the solar disk.



2-Day Conditions

Solar activity is expected to be at low levels with background X-ray flux at C-class levels.

Low wind speed and quiet geomagnetic conditions are expected over the weekend.

Enhanced HF conditions are expected for the next 2 days. It is advised to use the frequency ranges mentioned in the ionospheric section.

For information on radio blackout levels, please follow the link:

http://www.swpc.noaa.gov/noaa-scales-explanation

Acknowledgements:

<u>Images source</u>: Solar Dynamics Observatory-SDO) both images showing the Solar disk and Coronal Holes have been processed at SUPARCO using Automatic Solar Synoptic Analyzer (ASSA), developed jointly by the Korean Space Weather Centre of the Radio Research Agency (RRA) & Space Environment Laboratory (SELab).

<u>Data sources</u>: The planetary indices and solar data are taken from the URLs below:

http://www.spaceweather.go.kr http://www.sws.bom.gov.au http://www.solarmonitor.org

Sonmiani (SON): 25.2º N, 66.75º E Islamabad (ISB): 33.7º N, 73.13º E

ANNEXURE

	DEFINITIONS OF TERMINOLOGIES USED IN THIS SUMMARY			
foF2	Maximum frequency of F2-layer of the ionosphere			
h′F2	Virtual height of the F2-layer			
MUF	Maximum usable frequency for 3000 km			
K-index	Local index defining geomagnetic conditions			
Declination	Planetary A index defining geomagnetic conditions, predicted value during geomagnetic unsettled Conditions			
F	Magnitude of the total geomagnetic field vector (unit in nano Teslas)			
SON, difference	Sonmiani Geomagnetic Observatory mean value, <u>difference limit</u> from night time value of quiet conditions: 25-30 nT, max: 260 nT			
ISP	Islamabad Geomagnetic Observatory mean value			
SN	Relative sunspot numbers			
Vsw	Solar Wind Speed (km/s)			
F10.7	Solar radio flux at 2.8 GHz (10.7 cm wavelength)			
sfu	Solar flux unit (defines the solar radio 10.7 cm flux)			
Solar Flare	Could be B, C, M and X depending upon the intensity of x-rays being emitted (each type has further 10 classes based on amount of energy released by the flare)			
IMF	Interplanetary magnetic field (the source of which is the Sun)			
Bt	Total IMF (unit in Nano Teslas)			
Bz	Vertical component of IMF (could be north/upward/positive or south/downward/negative) (unit in nano Teslas)			
AR	Active Regions on the sun currently in view			
CME	Coronal Mass Ejection			
СН	Coronal Hole			
KASI	Korean Astronomy & Space Science Institute			
SWFs	Short-wave fadeouts, caused by M/X class flares on the daylit side of the hemisphere absorbing lower Frequencies and hampering HF communication.			
SSN-predicted	Smooth Sunspot Number-it is an estimated value using a mathematical relation to forecast it.			