

### 3-day Space Weather Conditions (SUPARCO)

Friday, March 26, 2021, 12:00 PST



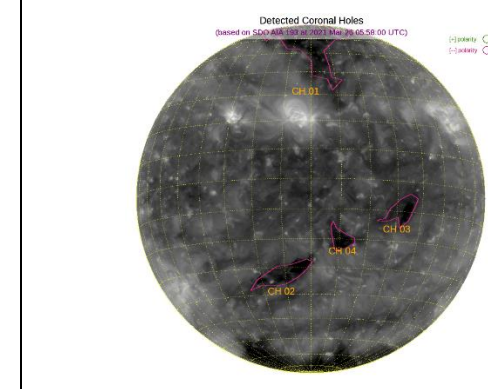
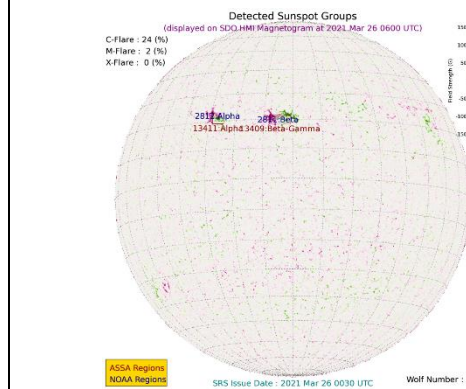
LOCAL CURRENT IONOSPHERIC CONDITIONS OVER PAKISTAN <sup>a</sup>								
DATE	26-Mar-21 (noon)			27-Mar-21 (noon)			28-Mar-21 (noon)	
foF2	9.2 MHz			7.9 MHz			9.4 MHz	
h'F2	298 km			290 km			260 km	
TEC	20 TECU			18 TECU			21 TECU	
Maximum Usable Frequency (MUF) and Optimum Traffic Frequency (FOT) for various distances								
Distance (km)	100	200	400	600	800	1000	1500	3000
MUF (MHz) for 3 days (26, 27, and 28 Mar)	9.3	9.7	11.0	12.9	15.0	17.1	22.1	30.3
	8.0	8.4	9.6	11.2	13.1	15.0	19.4	26.4
	9.6	10.1	11.8	14.1	16.7	19.2	25.0	33.4
MUF (MHz) for 3 days (26, 27, and 28 Mar)	7.9	8.2	9.4	11.0	12.7	14.6	18.8	25.7
	6.8	7.1	8.1	9.5	11.1	12.7	16.5	22.2
	8.1	8.6	10.0	12.0	14.2	16.4	21.2	28.4
Local ionospheric conditions are nominal with better MUF availability.								
LOCAL GEOMAGNETIC CONDITIONS OVER PAKISTAN <sup>ab</sup>								
K-index	2			Quiet to unsettled geomagnetic conditions expected			Quiet to unsettled geomagnetic conditions expected	
F	45455/50020 nT			45450±20 /50020±20 nT			45450±20/50020±20 nT	
The local geomagnetic field is quiet at the moment.								
SOLAR CONDITIONS								
SN	24			20 (SSN-predicted)			13 (SSN-predicted)	
F 10.7	79 sfu			80 sfu			75 sfu	
V <sub>sw</sub>	447 km/sec (varied in the past 12 hrs between 374 & 490 km/s)			Moderate solar wind expected			Elevated solar wind expected	
Solar flares	A 4.1(max. flare in the past 24 hrs: B 8.0 1519 UT Mar 25)			Low levels of solar activity expected			Low levels of solar activity expected	
IMF Bt	4.3 nT (varied in the past 12 hrs between 3.4 & 5.3 nT)			Expected to vary between positive and negative sectors.			Expected to vary between positive and negative sectors	
Bz	-3.6 nT (varied in the past 12 hrs between -3.6 & +0.9 nT)							

Solar conditions are at very low levels with background X-ray flux at A-class levels. Local HF working frequencies are normal as compared to monthly average predicted values.

### Daily Sun:26 March 2021

There are no active regions on the Sun capable of producing strong solar flares.

04 Coronal Holes (CHs) are detected on the solar disk.



### 2-Day Conditions

Solar activity is expected to be at very low levels with a slight chance for C-class flares.

Moderate to strong solar wind speed and quiet to unsettled geomagnetic conditions are expected due to the influence of 22 March CME and coronal hole effects over the weekend.

Normal HF conditions are expected for the next 2 days. It is advised to use the frequency ranges mentioned in the ionospheric section

For information on radio blackout levels, please follow the link:

<http://www.swpc.noaa.gov/noaa-scales-explanation>

### Acknowledgements:

Images source: Solar Dynamics Observatory-SDO) Both images showing the Solar disk and Coronal Holes have been processed at SUPARCO using Automatic Solar Synoptic Analyzer (ASSA), developed jointly by the Korean Space Weather Centre of the Radio Research Agency (RRA) & Space Environment Laboratory (SELab).

Data sources: The planetary indices and solar data are taken from the URLs below:

<http://www.spaceweather.go.kr>

<http://www.sws.bom.gov.au>

<http://www.solarmonitor.org>

<sup>a</sup>Sonmiani (SON): 25.2° N, 66.75° E

<sup>b</sup>Islamabad (ISB): 33.7° N, 73.13° E

ANNEXURE

<b>DEFINITIONS OF TERMINOLOGIES USED IN THIS SUMMARY</b>	
foF2	Maximum frequency of F2-layer of the ionosphere
h'F2	Virtual height of the F2-layer
MUF	Maximum usable frequency for 3000 km
K-index	Local index defining geomagnetic conditions
Declination	Planetary A index defining geomagnetic conditions, predicted value during geomagnetic unsettled conditions
F	Magnitude of the total geomagnetic field vector (unit in nano Teslas)
SON, difference	Sonmiani Geomagnetic Observatory mean value, <u>difference limit</u> from night time value of quiet conditions: 25-30 nT, max: 260 nT
ISP	Islamabad Geomagnetic Observatory mean value
SN	Relative sunspot numbers
Vsw	Solar Wind Speed (km/s)
F10.7	Solar radio flux at 2.8 GHz (10.7 cm wavelength)
sfu	Solar flux unit (defines the solar radio 10.7 cm flux )
Solar Flare	Could be B, C, M and X depending upon the intensity of x-rays being emitted (each type has further 10 classes based on amount of energy released by the flare)
IMF	Interplanetary magnetic field (the source of which is the Sun)
Bt	Total IMF (unit in nano Teslas)
Bz	Vertical component of IMF (could be north/upward/positive or south/downward/negative) (unit in nano Teslas)
AR	Active Regions on the sun currently in view
CME	Coronal Mass Ejection
CH	Coronal Hole
KASI	Korean Astronomy & Space Science Institute
SWFs	Short-wave fadeouts, caused by M/X class flares on the daylit side of the hemisphere absorbing lower Frequencies and hampering HF communication.
SSN-predicted	Smooth Sunspot Number-it is an estimated value using a mathematical relation to forecast it.