

3-day Space Weather Conditions (SUPARCO)

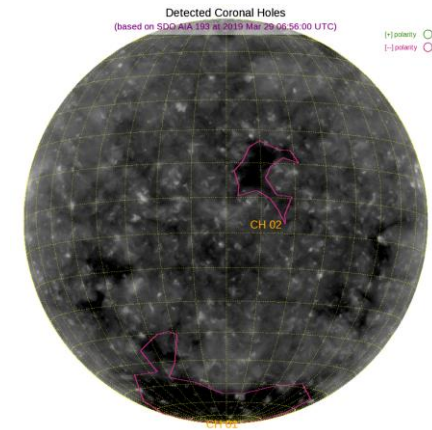
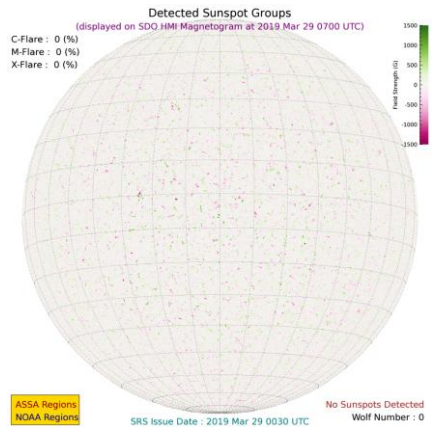
Friday, March 29, 2019, 15:08 PST



LOCAL CURRENT IONOSPHERIC CONDITIONS OVER PAKISTAN ^a																								
DATE	29-Mar-19 (noon)						30-Mar-19 (noon)						31-Mar-19 (noon)											
foF2	8 MHz						9 MHz						10 MHz											
h'F2	268 km						265 km						260 km											
TEC	13 TECU						15 TECU						18 TECU											
Maximum Usable Frequency (MUF) and Optimum Traffic Frequency (FOT) for various distances																								
Distance (Km)	100			200			400			600			800			1000			1500			3000		
MUF (MHz) for 3 days (29, 30, & 31 Mar)	8.1	9.2	10.2	8.5	9.6	10.7	9.9	11.2	12.6	11.8	13.4	15.0	13.9	15.8	17.8	16.0	18.2	20.5	20.8	23.6	26.6	28.0	31.7	35.5
FOT (MHz) for 3 days (29, 30, & 31 Mar)	6.9	7.8	8.7	7.3	8.2	9.1	8.4	9.5	10.7	10.1	11.4	12.8	11.8	13.4	15.1	13.6	15.5	17.4	17.7	20.0	22.6	23.8	26.9	30.2
Local ionospheric conditions are nominal with normal MUF conditions. It is advised to use the frequency range mentioned above w.r.t the distances (km).																								
LOCAL GEOMAGNETIC CONDITIONS OVER PAKISTAN ^{ab}																								
K-index	2						2						2											
F	45315 /49865 nT						45300±30 /49850±25 nT						45305±30 /49855±20 nT											
D	0.55 /0.73 degrees						0.55 /0.73 degrees						0.60 /0.72 degrees											
The local geomagnetic field is quiet at the moment.																								
SOLAR CONDITIONS																								
SN	0						2 (SSN-predicted)						2 (SSN-predicted)											
F 10.7	68 sfu						68 sfu						68 sfu											
V _{sw}	476 km/sec (varied in the past 12 hrs between 476 & 538 km/s)						Moderate solar wind speed expected						Solar wind speed expected to reach at ambient levels											
Solar flares	A 0.0 (max. flare in the past 24 hrs: A 0.0 0803 UT Mar 29)						Very low levels of solar flare activity expected with slight chances of C-class flare(s)						Very low levels of solar flare activity expected with slight chances of C-class flare(s)											
IMF Bt	4.7 nT (varied in the past 12 hrs between 3.6 & 5.8 nT)						Expected to remain.						Expected to remain positively oriented but might switch to negative sector.											
Bz	+0.8 nT (varied in the past 12 hrs between -3.2 & +0.8 nT)																							
Solar conditions are at <u>very low</u> levels with background X-ray flux at A-class levels. Local HF working frequencies are <u>normal</u> as compared to monthly average predicted values.																								
Daily Sun: 29 March 2019																								

Solar disk is currently spotless therefore; there are no chances of strong solar flare(s) to erupt.

02 Coronal Holes (CH) are detected on the solar disk, both with negative polarity which might keep the solar wind elevated.



2-Day Conditions

Solar activity is expected to remain at very low levels over the weekend.

Quiet to slightly unsettled geomagnetic conditions are expected due to high speed solar wind streams associated with current coronal holes and CIR effects. The solar wind speed is expected to return to ambient levels as the CIR and CH high speed streams wane in the next two days.

Nominal HF conditions are expected for the weekend. It is advised to use the frequency ranges mentioned in the ionospheric section.

For information on radio blackout levels, please follow the link:

<http://www.swpc.noaa.gov/noaa-scales-explanation>

Acknowledgements:

Images source: Solar Dynamics Observatory-SDO) Both images showing the Solar disk and Coronal Holes have been processed at SUPARCO using Automatic Solar Synoptic Analyzer (ASSA), developed jointly by the Korean Space Weather Centre of the Radio Research Agency (RRA) & Space Environment Laboratory (SELab).

Data sources: The planetary indices and solar data are taken from the URLs below:

<http://www.spaceweather.go.kr>

<http://www.sws.bom.gov.au>

<http://www.solarmonitor.org>

^aSonmiani (SON): 25.2° N, 66.75° E

^bIslamabad (ISB): 33.7° N, 73.13° E

ANNEXURE

DEFINITIONS OF TERMINOLOGIES USED IN THIS SUMMARY	
foF2	Maximum frequency of F2-layer of the ionosphere
h'F2	Virtual height of the F2-layer
MUF	Maximum usable frequency for 3000 km
K-index	Local index defining geomagnetic conditions
Declination	Planetary A index defining geomagnetic conditions, predicted value during geomagnetic unsettled conditions
F	Magnitude of the total geomagnetic field vector (unit in nano Teslas)
SON, difference	Sonmiani Geomagnetic Observatory mean value, <u>difference limit</u> from night time value of quiet conditions: 25-30 nT, max: 260 nT
ISP	Islamabad Geomagnetic Observatory mean value
SN	Relative sunspot numbers
V _{sw}	Solar Wind Speed (km/s)
F10.7	Solar radio flux at 2.8 GHz (10.7 cm wavelength)
sfu	Solar flux unit (defines the solar radio 10.7 cm flux)
Solar Flare	Could be B, C, M and X depending upon the intensity of x-rays being emitted (each type has further 10 classes based on amount of energy released by the flare)
IMF	Interplanetary magnetic field (the source of which is the Sun)
B _t	Total IMF (unit in nano Teslas)
B _z	Vertical component of IMF (could be north/upward/positive or south/downward/negative) (unit in nano Teslas)
AR	Active Regions on the sun currently in view
CME	Coronal Mass Ejection
CH	Coronal Hole
KASI	Korean Astronomy & Space Science Institute
SWFs	Short-wave fadeouts, caused by M/X class flares on the daylit side of the hemisphere absorbing lower frequencies and hampering HF communication.
SSN-Predicted	Smooth Sunspot Number-it is an estimated value using a mathematical relation to forecast it.
CIR	Co-rotating Interaction Region